

# ATLAS OF PECULIAR GALAXIES

HALTON ARP

Mount Wilson and Palomar Observatories

Carnegie Institution of Washington, California Institute of Technology

*Received December 27, 1965; revised April 4, 1966*

## ABSTRACT

The *Atlas of Peculiar Galaxies* presents the results of more than four years of direct photography with the 200-inch telescope. Unusual galaxies were selected from lists by Zwicky, Vorontsov-Velyaminov, and unpublished lists by A. G. Wilson, E. Herzog, Wirtanen, the author, and others. Plate files of Mount Wilson and Palomar Observatories were searched for suitable objects, and some of these plates are reproduced. For the most part, however, limiting, good-seeing exposures were obtained on blue-visual sensitive plates at the prime focus ( $11''/mm$ ) of the 200-inch.

The objects are arranged 6 on a page, 57 pages, for a total of 338 objects. Magnifications range from  $1\times$  to  $10\times$ . They are ordered empirically according to their form and visual appearance. Tabular material presents known data for the objects including positions, plate data, radial velocities when known, references, and remarks on both over-all and detailed peculiarities.

The *Atlas* is also available in large size,  $11 \times 14$ -inch photographic reproduction, from the California Institute of Technology Bookstore for a price of about \$60 bound.

## PREFACE

Forty years after the discovery that galaxies were independent stellar systems, we still have not penetrated very far into the mystery of how they maintain themselves or what physical forces are responsible for shaping their observed forms. The galaxies are the constituent units of mass and energy in the Universe, and yet we are still challenged by such questions as: What causes the characteristic shape of spiral galaxies? How are elliptical galaxies related to spirals? How are galaxies formed, and how do they evolve?

It is difficult to resist an oversimplified impression of what a galaxy is because the Hubble classification divides the galaxies into the well-known categories of smooth, amorphous ellipticals, and flattened spirals with star-studded arms. But far from all galaxies fit the Hubble sequence of nebular forms. In fact, when looked at closely enough, every galaxy is peculiar. Appreciation of these peculiarities is important in order to build a realistic picture of what galaxies are really like. But the peculiarities are also important for the reason that, if we could analyze a galaxy in the laboratory, we would deform it, shock it, probe it in order to discover its properties. The peculiarities of the galaxies pictured in this *Atlas* represent perturbations, deformations, and interactions which should enable us to analyze the nature of the real galaxies which we observe and which are too remote to experiment on directly. In general, the more conspicuous the peculiarity, the more illustrative it is of special events and reactions that occur in galaxies. From this range of experiments which nature furnishes us, then, it is our task to select and study which will give the most insight into the composition and structure and the forces which govern a galaxy.

The present *Atlas* specifically started from an attempt to better understand spiral galaxies. Despite even recent analyses from a contrary standpoint, I believe that gravitational orbits in a stellar assemblage will not alone furnish satisfactory explanations of galaxies. It is clear that the convolution which spiral arms are seen to undergo in certain galaxies cannot be performed by loci of stellar orbits. In the investigation of these special spiral properties, therefore, galaxies which showed unusual or perturbed arms or filamentary extensions were sampled with high-resolution photographs with the Palomar 200-inch telescope. Subjects were first drawn from the pioneering work of Zwicky and Vorontsov-Velyaminov. So many important objects emerged under high-resolution,

limiting-magnitude study, however, that the investigation into the nature of spiral arms was temporarily postponed in order to organize systematically these new phenomena into groups and publish a representative sample of the best objects.

The *Atlas* as it has been realized in the following pages illustrates again that galaxies cannot be characterized as just assemblages of stars, radiation, and gravitation. The following *Atlas* pictures emphasize the importance of dust in some; they particularly imply a much more important role for the gas in general and point to the existence of either new forces or forces which previously have been little considered. For example, the twisted, distorted shapes and curious linkages pictured here attest to the fact that there are viscosity-like forces present that in some cases are dominant. Probably these forces are due to magnetic effects. Vorontsov-Velyaminov has stressed in the past the probable magnetic nature of these effects. Magnetic forces are very difficult to study, but may be very important in our Universe. The recent radio-astronomy discoveries of violent events in galaxies reveal sources of energetic charged particles. These charged particles interact with magnetic fields and offer the hope of mapping, measuring, and understanding cosmic magnetic fields. Exploration of the connection between the plasmas observed with the radio telescopes and the optical evidences of plasma effects pictured in the present *Atlas* is now open to us.

The over-all aim of this *Atlas* is to present a number of examples of various kinds of peculiar galaxies. They are presented in groupings that appear roughly similar, thereby furnishing also a rough, initial classification. Phenomena which each group represent may then be investigated by picking the most favorable members in size or brightness, studying different members of the group in different orientations, and, finally, making some preliminary statistics of certain kinds of phenomena and their relationship to other observable parameters. It is hoped that this investigative procedure will not only clarify the workings of galaxies themselves but will also reveal physical processes and how they operate in galaxies, and ultimately furnish a better understanding of the workings of the Universe as a whole.

It is a pleasure to acknowledge the help of William Miller, who photographically copied the original glass negatives; Lowell Peterson, of Graphic Arts at the California Institute of Technology, who supervised the large-size photographic reproduction of the *Atlas*; Frank Brueckel, who carried out many computational tasks connected with the *Atlas*; and, of course, all those astronomers who suggested candidates for the *Atlas* from their own personal store of knowledge and who gave advice and encouragement.

#### INTRODUCTION

The *National Geographic Society-Palomar Observatory Sky Survey* was completed in 1956. For seven years the 48-inch Schmidt telescope had surveyed the sky north of  $\delta = -27^\circ$ . The 1758 highest-quality plates that were finally accepted penetrated about three times deeper into space than any previous survey had ever reached. Astronomers are still studying and cataloguing the information contained in this survey, and will continue to do so for many years to come.

One of the first astronomers to use the prints of the *Sky Survey Atlas* for a systematic study was Professor Vorontsov-Velyaminov of the Sternberg Astronomical Institute in Moscow (2).<sup>1</sup> In 1959 he published positions, with copies of *Sky Survey* pictures, of 355 peculiar and interacting galaxies that he had discovered on *Survey* prints. The publication of this list enabled the undertaking of one kind of project for which the 48- and 200-inch telescopes on Palomar Mountain were originally designed. The fast-focal-ratio, wide-field Schmidt telescope was intended to survey objects of interest. The maximum light-gathering power and resolution of the 200-inch could then be turned individually on the most interesting objects.

<sup>1</sup> For numbered references see the Bibliography at the end of the text.

## ATLAS OF PECULIAR GALAXIES

## 3

When selected members of Vorontsov-Velyaminov's catalogue were photographed with the 200-inch, some turned out to be much more interesting than on the smaller-scale plates, while others turned out to be less interesting or ordinary. After some preliminary experience with the 200-inch scale, it soon became possible to inspect the Vorontsov-Velyaminov objects first on the *Survey* prints to cull out the less interesting objects. In the process of inspecting these objects and checking their positions, other very unusual galaxies were noticed on the same *Survey* prints and included in the 200-inch program. This demonstrated that not all the important objects had been catalogued, and efforts were made to compile from other sources a more complete list of candidates for peculiar galaxies.

One additional source of peculiar-galaxy candidates was the set of notes which A. G. Wilson had made upon inspecting the original *Sky Survey* plates as they were taken. These were kindly put at my disposal. Another list of peculiar objects was given me by E. Herzog, who has carefully searched the *Survey* plates for such objects. Thornton Page contributed a list of peculiar objects he knew and a list of peculiar galaxies which C. A. Wirtanen had compiled from the *Lick Position Survey*. Holmberg's pairs of galaxies were inspected.

Special objects were also contributed by W. W. Morgan, F. Zwicky, Charles Kowal, and Gibson Reaves. Finally, the plates of Minkowski and Baade, which are stored at the Mount Wilson Observatory, were searched for peculiar objects. A surprising result was that none of these lists, including my own, had very much overlap with one other. The conclusion seems to be that, aside from the brighter and therefore well-known peculiar galaxies, the fainter peculiars have not been fully catalogued, and that the fainter peculiar galaxies pictured in this *Atlas* represent only a sample of that group.

At first the photographs with the 200-inch were made with various plate and filter combinations to discover in which wavelengths the peculiar features would show best. Although red wavelengths sometimes showed features better, in general, the filaments, connections, and faint outer features were more conspicuous on blue-sensitive (Eastman Kodak 103aO) plates. At that time, however, the sky was becoming so dark because of sunspot minimum, that it was possible to reach fainter limiting magnitudes by exposing blue plates for 60–70 min. To make the project possible in terms of available observing time, the band pass was widened by using 103aD plates and including the visual as well as blue wavelengths in a limiting exposure of the order of 30 min. Finally, it became clear that the night-sky emission line at  $\lambda$  5577 was contributing appreciably to the brightness of night-sky background, and the emulsion was changed to 103aJ from then until the conclusion of the project. The 103aJ plates registered light roughly between the  $\lambda$  3600 cutoff of the f/3.67 corrector lens on the 200-inch telescope and the  $\lambda$  5400 photographic emulsion cutoff. That, in general, is the region of maximum contrast for galaxies (10), and the very deep exposures made here (to densities of 0.7 to 1.0 for sky background), the very dark night skies, and the 20 per cent increases in development time give, on the average, a set of photographs that show fainter stars—and particularly fainter surface-brightness features—than ever before detected in galaxy subjects. The reproduction of these prints in the Mount Wilson and Palomar photographic laboratory by William Miller was a difficult job which was carefully controlled so that almost all the original features on the plates, even the faintest, are reproduced in the *Atlas*.

Whenever possible, poor-seeing plates were repeated under better seeing conditions, so that the final *Atlas* contains only plates taken with seeing 2 or better. The star images on the plates taken with the 200-inch presented in this *Atlas* are therefore generally between 1" and 2" diameter. Search of the Observatory plate records located some of the prospective *Atlas* galaxies which had been already photographed. I am grateful to Zwicky, Sandage, and Baum for allowing me to reproduce some of the photographs of these objects, and they are credited under the listed plate numbers in Table 1. Most of the 338 photographs shown in the *Atlas* are from plates taken with the 200-inch tele-

scope. Occasionally a very large object is shown in a print from a 48-inch telescope plate (designated PS) in order to emphasize its correct sequence in the order of forms.

Because so many of the physical processes pictured are not understood, no rigorous attempt at classification has been made. The galaxies have been grouped empirically, putting together all the objects that look alike. Special emphasis is on the form of the galaxies or the nature of the peculiarity and the gradual change of the peculiarity from object to object. Sometimes an object will belong in more than one category, and then it is cross-referenced in Table 1 or shown under different magnification in different sections of the *Atlas*. The schematic plan of arrangement of the different kinds of galaxies is shown in Figure 1. The largest class involves peculiar spiral galaxies (Nos. 1–102). The largest subclass of peculiar spirals are spirals with companions attached to spiral arms (Nos. 37–102). Then there is a group of elliptical or E-like galaxies (Nos. 102–145). Of course, there is overlap, and in the very interesting group ranging from Nos. 91 to 114 it is impossible to say whether the E is a companion to the spiral galaxy or vice versa. The third major group (Nos. 146–268) involves galaxies or groups of objects that are not primarily classifiable as either E's or spirals, or whose most outstanding peculiarity does not fall in the first two major categories. In the fourth major category (Nos. 269–327), group character is the most important consideration. Six objects classifiable only as miscellaneous are shown at the end (Nos. 332–338).

When possible, information has been gathered from the literature regarding apparent magnitude, redshift velocities, and any known spectral peculiarities. Table 2 lists all the objects in this *Atlas* in order of right ascension and gives references to known redshift velocities. In Table 3 all coincidences of *Atlas* objects with catalogue radio sources are noted and referenced. With the exception of bright radio sources such as Fornax A, *Atlas* objects were not selected because they were radio sources—although Minkowski's plates were generally taken in search of radio-source identifications. In many cases, however, nothing more is known about an object than what is shown in the *Atlas*. An important task in the future will be to undertake photometric and spectroscopic observations of these objects. Then, when distances, absolute magnitudes, and spectral characteristics are known, a more meaningful classification and interpretation of the objects in this *Atlas* can take place.

#### THE ATLAS AND THE CATALOGUE

The 338 photographs shown in the following fifty-seven pages of the *Atlas* all have a notch marking the north point. West is  $90^\circ$  clockwise. The prints represent magnifications from the original plates of  $1\times$ ,  $2\times$ ,  $4\times$ ,  $6\times$ ,  $8\times$ , and  $10\times$ . Since all the 200-inch plates in this program were taken with the Ross f/3.67 corrector lens, the scales on the original prints therefore vary from  $11''/\text{mm}$  to  $1.^{\circ}1/\text{mm}$ . The natural scale of the few prints from Schmidt plates is  $67''/\text{mm}$ . In reproduction of the large-size photographic edition, all these scales have been reduced by a factor of 0.97. In the reproduction in the *Astrophysical Journal Supplement*, the original print scale has been reduced by a factor of 0.54.

About one-third of the prints were made by an automatic, fluorescent screen dodging process, i.e., by compressing the density range so that one can see very faint features and yet see into the brighter inner regions on the same print. In some cases the automatic dodging has introduced slightly lighter halos around the stars.

#### THE CATALOGUE

*Col. 1:* Identification number in this catalogue. See Figure 1 for arrangement of types of objects.

*Cols. 2–3:* Right ascension and declination of objects for 1970 epoch. Positions are from three sources: (1) NGC positions where available. If more than one NGC object is pictured, the position of the westernmost (smallest number) is given. (2) Positions from

## ATLAS OF PECULIAR GALAXIES

## 5

200-inch dial readings calibrated by objects with known positions. (3) Measurements on 48-inch *Sky Survey* plates (whenever possible, differential measurements from nearby NGC objects). The final accuracy of these positions, from cross-checking the different methods is on the average better than  $\pm 0^{\text{m}}2$  in R.A. and  $\pm 2'$  in decl. A few positions are from Vorontsov-Velyaminov.

*Col. 4:* Designation. NGC or IC numbers are given when object has one; otherwise designation is blank.

*Col. 5:* Plate number. PH designates 200-inch Hale telescope; PS designates 48-inch Schmidt. Plates taken by Arp unless designated B = Baade, Bm = Baum, M = Minkowski, S = Sandage, Z = Zwicky.

*Col. 6:* Exposure in minutes.

*Col. 7:* Kind of emulsion used—"bk" designates baked and "b" designates lightly baked; "exp" designates experimental.

*Col. 8:* Identifies the filter used, if any.

*Col. 9:* Seeing. On a scale in which image size is about 1" to  $1\frac{1}{2}$ " for seeing 3. Each unit poorer than three-image size roughly doubles and approximately halves for each unit better.

*Col. 10:* Magnification (varies from  $1\times$  to  $10\times$ ).

*Col. 11:* Source. As far as can be determined, the person who first noticed the peculiar object is named. Vorontsov-Velyaminov ("VV") numbers are given when they exist. "DDO" is David Dunlap Observatory.

*Col. 12:* Remarks on objects shown in photographs. Major peculiarities are described in Figure 1; additional peculiarities and remarks are noted here.

## BIBLIOGRAPHY AND REFERENCES

[NOTE.—The numbered references below are cited by number in Tables 1 and 2. There is no Reference No. (42). The references in categories A and B are in chronological order; those in category C are in alphabetical order except No. (47), which was added to the list later.]

## A. GENERAL CATALOGUES AND COMPILATIONS OF GALAXY OBSERVATIONS

Dryer, J. L. E. 1888, *Mem. R. Astro. Soc.*, 49 (*New General Catalogue of Nebulae and Star Clusters*), Index Cat., Vols. 51 and 59.

Shapley, H., and Ames, A. 1932, *Harvard Ann.*, 88, No. 2.

Holmberg, E. 1937, *Lund Ann.*, 6 (*Catalogue of Double and Multiple Galaxies*).

- (1) Humason, M. L., Mayall, N. U., and Sandage, A. R. 1956, *A.J.*, 61, 97.  
Bergh, S. van den. 1959, *Pub. David Dunlap Obs.*, 2, 147.
- (2) Vorontsov-Velyaminov, B. A. 1959, *Atlas and Catalogue of Interacting Galaxies*, Vol. 1 (Sternberg Institute, Moscow State University, Moscow).
- (3) Sandage, A. R. 1961, *Hubble Atlas of Galaxies* (Carnegie Institution of Washington, Washington, D.C.)  
Zwicky, F., Herzog, E., and Wild, P. 1961, *Catalogue of Galaxies and Clusters of Galaxies*, Vol. 1 (California Institute of Technology, Pasadena).  
Burbidge, E. M., and Burbidge, G. R. 1962, *A.p. J.*, 135, 694.  
Vorontsov-Velyaminov, B. A., and Krasnogorskaya, A. A. 1962, *Morphological Catalogue of Galaxies*, Part I (Moscow State University, Moscow).  
Zwicky, F., and Herzog, E. 1963, *Catalogue of Galaxies and Clusters of Galaxies*, Vol. 11 (California Institute of Technology, Pasadena).
- (4) Vaucouleurs, G. de., and Vaucouleurs, A. de. 1964, *Reference Catalogue of Bright Galaxies* (University of Texas Press, Austin).  
Vorontsov-Velyaminov, B. A., and Arhipova, V. P. 1964, *Morphological Catalogue of Galaxies*, Part II (Moscow State University, Moscow).

## B. CONFERENCE PROCEEDINGS

*Proceedings of the Symposium on the Motion of Gaseous Masses of Cosmical Dimensions, "Problems of Cosmical Aerodynamics," Paris, August, 1949* [Central Air Documents Office, Dayton, Ohio, 1951 (AF-WPAFB-O-25)].

- Comparison of the Large-Scale Structure of the Galactic System with That of Other Stellar Systems*, I.A.U. Symposium No. 5, ed. N. G. Roman (Cambridge: Cambridge University Press, 1958).
- Les Recherches Galactiques et Extra Galactiques et la Photographie Électronique*, Paris, June-July 1959, *Ann. d'ap.*, 23, 305-366, 1960.
- Conference on the Instability of Systems of Galaxies*, Santa Barbara, California, August 8-9, 1961, *A.J.*, 66, 533-636, 1961.

## C. OBSERVATIONS OF SPECIFIC GALAXIES

- (5) Ambartsumian, V. A. 1961, *A.J.*, 66, 537.
- (6) Arp, H. C. 1962, *Ap.J.*, 136, 1148.
- (7) ———. 1964, *ibid.*, 139, 1378.
- (8) ———. 1965, *Science*, 148, 363.
- (9) Baade, W. 1956, *Ap.J.*, 123, 550.
- (10) Babcock, H. W., and Johnson, J. J. 1941, *Ap.J.*, 94, 271.
- (11) Burbidge, E. M., and Burbidge, G. R. 1959a, *Ap.J.*, 129, 271.
- (12) ———. 1959b, *ibid.*, 130, 12.
- (13) ———. 1959c, *ibid.* p. 15.
- (14) ———. 1959d, *ibid.*, p. 20.
- (15) ———. 1959e, *ibid.*, p. 23.
- (16) ———. 1959f, *ibid.*, p. 26.
- (17) ———. 1961a, *ibid.*, 133, 726.
- (18) ———. 1961b, *A.J.*, 66, 541.
- (19) Burbidge, E. M., Burbidge, G. R., and Hoyle, F. 1963, *Ap.J.*, 138, 873.
- (20) Burbidge, E. R., Burbidge, E. M., and Sandage, A. R. 1963, *Rev. Mod. Phys.*, 35, 947.
- (21) Burbidge, E. M., Burbidge, G. R., and Rubin, V. C. 1964, *Ap.J.*, 140, 942.
- (22) Burbidge, E. M., and Burbidge, G. R. 1964a, *Ap.J.*, 140, 1307.
- (23) Burbidge, E. M. 1964b, *Ap.J.*, 140, 1617.
- (24) Burbidge, E. M., Burbidge, G. R., and Prendergast, K. H. 1964, *Ap.J.*, 140, 1620.
- (25) Burbidge, E. M., and Burbidge, G. R. 1965, *Ap.J.*, 142, 634.
- See also reference added in proof (47).
- (26) Demoulin, M. 1965, *Compte rend.*, 260, 3287.
- (27) Haro, G. 1956, *Bol. Obs. Tonantzintla y Tacubaya*, No. 14, p. 8.
- (28) Haro, G., and Münch, G. 1958, *Sky and Telescope*, p. 231 (March).
- (29) Lynds, C. R., and Sandage, A. R. 1963, *Ap.J.*, 137, 1005.
- (30) Markarian, B. E. 1961, *A.J.*, 66, 555.
- (31) Mayall, N. U. 1948, *Pub. A.S.P.*, 60, 266.
- (32) ———. 1958, *Large-Scale Structure of the Galactic System*, I.A.U. Symp. No. 5, ed. N. G. Roman, p. 3; *Lick Obs. Sec. II*, No. 81.
- (33) Morgan, W. W., and Mayall, N. U. 1957, *Pub. A.S.P.*, 69, 291.
- (34) Osterbrock, D. E. 1960, *Ap.J.*, 132, 325.
- (35) Page, T. 1952, *Ap.J.*, 116, 63.
- (36) Sandage, A. R. 1963, *Ap.J.*, 138, 863.
- (37) Sandage, A. R., and Miller, W. C. 1964, *Science*, 144, 405.
- (38) Searle, L. 1965, *Nature*, 207, 1282.
- (39) Seyfert, C. K. 1943, *Ap.J.*, 97, 28.
- (40) Vaucouleurs, G. de, and Vaucouleurs, A. de. 1963, *Ap.J.*, 137, 363.
- (41) Wade, C. M. 1961, *Pub. NRAO* (Green Bank), 1, 99.
- (43) Zwicky, F. 1939, *Proc. Nat. Acad. Sci.*, 23, 251.
- (44) ———. 1956, *Ergebnisse der exakten Wissenschaften*, 29, 344.
- (45) Zwicky, F., and Humason, M. L. 1961, *Ap.J.*, 133, 794.
- (46) Zwicky, F. 1966, 1965 Palomar supernova search, *Pub. A.S.P.*, to be published.
- (47) Burbidge, E. M., and Burbidge, G. R. 1964, *Ap.J.*, 140, 1445.

TABLE 1  
DATA FOR ILLUSTRATIONS

No.	$\alpha$ (1970)	$\delta$ (1970)	Designation	Plate No.	Exp. (min)	Emul.	Fil- ter	See- ing	Mag	Source	Remarks
1	9 22.9	+49 28	NGC 2857	PH-4448	30	103a-J	-	2	4X	DDO 204	High contrast print of low surface brightness spiral.
2	16 15.1	+47 7		3990	30	103a-O 1b	GG 13	3	6X	DDO 214	Low surface brightness dwarf. Large bright knot in arm appears almost stellar.
3	22 34.8	- 3 6		3987	45	103a-O 1b	GG 13	3	4X	DDO 95, VV 251	Low surface brightness dwarf.
4	1 47.0	-12 31	NGC 3664	4014	30	103a-O 1b	GG 13	4	6X	DDO 14	Not known if both galaxies are at the same distance.
5	11 22.7	+ 3 29	NGC 2557	4139	25	103a-D	-	3	8X	VV138	Fainter dwarf 6.8.
6	8 11.1	+46 5	NGC 2557	4122	30	103a-D	-	3	8X	VV28	Bifurcated arm does not start at end of bar.
7	8 49.1	-16 30		4125	30	103a-D	-	2	6X		Nucleus off center in ring.
8	1 20.9	- 1 1	NGC 497	4387	25	103a-J	-	2	6X		Position of larger spiral. Outer arms do not start at termination of bar.
9	8 11.6	+73 42	NGC 2323	4689	25	103a-D	-	3	8X		Nucleus may be double or superposed star.
10	2 16.6	+ 5 30		4360	25	103a-D	-	3	4X		High surface brightness.
11	1 7.9	+14 11		4278	30	103a-D	-	3	4X		Almost no nucleus. (3)
12	8 33.4	+28 41	NGC 2608	4348	40	103a-O	-	1-2	6X		Feature appears to be a ruptured or obscured ring. Member of group. (18)
13	22 58.6	+15 49	NGC 7448	3999	30	103a-O 1b	GG 13	4	8X		See also 317. Large concentration at end of S arm. (3)
14	22 34.1	-26 12	NGC 7314	3994	30	103a-O 1b	GG 13	3	4X		End of one spiral arm partially disconnected. (3) (33)
15	22 50.0	- 6 43	NGC 7393	4011	30	103a-D 1b	-	4	8X		
16	11 18.6	+13 11	NGC 3627	3540Z	30	103a-D	GG 11	2-3	2X		
17	7 41.3	+73 52		4378	25	103a-D	-	3	10X	VV349	
18	12 4.0	+50 42	NGC 4088	4426	30	103a-J	-	1-2	4X		
19	0 30.2	- 5 19	NGC 145	266M	20	103a-O	-	1-2	6X		
20	4 18.4	+ 2 1		276M	20	103a-O	-	1-2	10X		
21	11 3.2	+30 15		4451	30	103a-J	-	2	10X		
22	11 58.0	-19 6	NGC 4027	4454	25	103a-J	-	2	2X	VV66	
23	12 40.1	+41 19	NGC 4618	4456	30	103a-J	-	3	1X	VV73	
24	10 52.8	+57 9	NGC 3445	4393	25	103a-D	-	2	6X	VV14	
25	7 17.9	+85 50	NGC 2276	4418	30	103a-D	GG 11	2	4X		
26	14 2.1	+54 29	NGC 5457	PS-8271	50	103a-J	Wr. #4	4-3	4X*	M101, VV344	
27	11 19.4	+53 21	NGC 3831	PH-4477	20	103a-J	-	3	4X		
28	23 27.1	+22 15	NGC 7678	4683	30	Ia-O bkd	-	3	8X		Note straight arms, absorption tube crossing from inside to outside of S arm.
29	20 34.3	+60 2	NGC 6946	3973	25	103a-O	-	3	1X		Note straight heavy arm.
30	17 22.6	+62 11		4275	30	103a-D	-	3	8X		Supernova once observed in tip of thick arm. (31)(46)
31	1 49.4	+21 45		4315	30	103a-O	-	2	6X		Comp. appears physically connected to flat-on spiral system.
32	17 12.5	+59 22		4274	30	103a-D	-	3	10X		High surface brightness irregularity is 5°N.
33	13 36.1	+ 6 35		4537	25	103a-J	-	3	6X		Position of larger spiral. See 326 for smaller scale picture. Part of galaxy chain.
34	12 40.0	+26 14		4479	30	103a-J	-	2-3	4X	VV257	
35	0 20.7	- 1 34		4371	40	103a-O	-	3	8X		Radio source M00-01 is 50°W.
36	13 32.9	+31 35		4458	30	103a-J	-	3	8X		Knots in arms approach appearance of small companions.
37	2 41.1	- 0 8	NGC 1068	4288	15	103a-O	Polaroid	2	6X		Seyfert galaxy. Small knot in arm. (3)(16)(39)
38	17 30.1	+75 45	NGC 6412	4691	30	103a-O	-	3	6X		Small ring in arm on N side, part of large ring on E side shows in Hα only.
39	3 28.5	-22 22	NGC 1347	4110	30	103a-D	-	1	6X		Absorption off edge of small galaxy obscure part of large galaxy.
40	13 28.1	+37 34	IC 4271	4185	30	103a-D	-	3	10X		Possible connection.
41	3 8.4	-20 41	NGC 1232	4308	30	103a-O	-	3	2X		Companion spiral wound in same sense as parent. Note split of companion's arm further into center. (3)

\* Plate taken with 48-inch Schmidt.

TABLE I (cont'd)

No.	$\alpha$ (1970)	$\delta$ (1970)	Designation	Plate No.	Exp. (min)	Fil- ter	See- ing	Mag.	Source	Remarks
42	15 1. <sup>b</sup> 4	+23 27	NGC 5839	PH-4255	30	103a-D	-	2	6X	VV77
43	10 22.3	+16 53		4421	35	103a-O	-	3	8X	
44	10 24.1	- 2 3	IC 609	4168	30	103a-D	-	2-3	6X	VV354
45	14 18.7	+52 0		4254	30	103a-D	-	2	8X	VV2
46	23 32.2	+29 52		4284	30	103a-D	-	2	10X	VV314
47	14 45.8	+18 59		4487	35	103a-J	-	3	8X	
48	1 18.5	+12 11		4374	25	103a-D	-	3	10X	
49	14 30.9	+ 8 12	NGC 5666	4528	25	103a-J	-	3	8X	
50	23 56.4	-14 12	IC 1520	4067	30	103a-D	-	1	10X	VV25
51	0 4.2	-13 36		4088	30	103a-D	-	3	10X	
52	6 18.0	+ 3 42		4010M	25	103a-O	-	2-3	10X	
53	10 33.6	-17 00	NGC 3290	4452	25	103a-J	-	2	8X	Wirtanen 18
54	2 23.0	- 4 47		4390	25	103a-D	-	2	8X	Wilson 14
55	9 13.9	+44 27		4163	30	103a-D	-	2	8X	VV155
56	1 56.6	+17 4		4117	30	103a-D	-	2	10X	VV12
57	13 16.2	+14 36		4482	25	103a-J	-	2	8X	VV298
58	8 30.2	+19 19		4362	25	103a-D	-	2-3	8X	Wilson 18
59	0 59.0	- 9 19	NGC 341	4343	30	103a-D	-	1	8X	Wilson 8
60	13 13.2	+26 16		4636	25	103a-J	-	3	10X	
61	4 35.2	- 2 21		284M	20	103a-O	-	1	10X	
62	11 52.1	+43 36		4495	25	103a-J	-	1-2	10X	VV286
63	9 37.4	+32 27	NGC 2944	4100	25	103a-D	-	3	10X	VV82
64	14 43.8	+19 36		4529	30	103a-J	-	3	8X	Herzog 1
65	0 20.3	+22 13		4358	30	103a-J	-	3	4X	Wilson 2
66	16 26.1	+51 36		4670	30	103a-O	-	2	8X	
67	1 19.8	- 0 42	NGC 7756	4359	25	103a-D	-	2	8X	
68	23 47.0	+ 3 57	NGC 5579+80	4678	30	103a-J	-	2	6X	
69	14 19.2	+95 19		4246	30	103a-D	-	3	8X	VV142
70	1 21.8	+30 37		4303	30	103a-D	-	3	8X	VV341
71	16 3.8	+17 46	NGC 5996+94	4273	30	103a-D	-	2	10X	VV16
72	15 46.7	+18 00		4253	30	103a-D	-	3	6X	
73	16 34.2	+46 17	IC 1222	4541	25	103a-J	-	3	8X	
74	2 6.4	+41 20		4361	25	103a-D	-	2-3	8X	Wilson 12
75	1 49.8	- 4 12	NGC 702	267M	20	103a-O	-	1	6X	
76	12 35.3	+13 19	NGC 4569	4181	30	103a-D	-	3	2X	M 90
77	2 45.2	-30 24	NGC 1097	4662	20	103a-J	-	3	2X	
78	1 57.7	+18 52	NGC 772	4305	45	103a-O	-	3	2X	
79	14 08.5	+17 46		4428	25	103a-D	-	1	10X	
80	8 45.0	+74 14	NGC 2639	4690	20	103a-J	-	2	6X	
81	18 13.4	+68 18	NGC 6621+22	3998	45	103a-D	-	3	8X	VV247
82	8 9.4	+25 18	NGC 2535+36	4085	30	103a-D	-	3	4X	VV9
83	11 38.6	+15 29	NGC 3799+3800	4422	25	103a-D	-	2-3	6X	VV350
84	13 57.4	+37 35	NGC 5334+95	4187	30	103a-D	-	3	4X	VV48
85	13 28.6	+47 21	NGC 5194+95	PS-5559	40	103a-D	#23A	3	6X*	M 51, VV1
86	23 45.6	+29 19	NGC 7733+52	PH-3986	45	103a-D	-	2-3	4X	VV5

Faint bifurcated arm to comp., one faint arm on comp. coll'd  
same direction as parent. (3)  
One side of ring obscured or disrupted; other side has low S. B.  
comp.

Directions inverted; N is opposite tab mark. One arm leads  
toward large comp.; other toward small comp.  
Companion connected to main spiral.

Appearance of wake from stellar object in E arm.

Radio source M10-17 is 1.<sup>b</sup>6 W, 10°S.  
Arm toward companion split, contains no hole.  
Arm has four separate condensations in line. North at bottom.

Small companion connected to end of arm.  
Comp. on end of broken arm nearly star-like.

Third arm in direction of companion.  
One arm leads to companion.  
High surface brightness companion.

Both arms lead toward companions.  
Both of open spiral. Comps. lie off projected ends of both  
spiral arms.

Comps. lie on inner and outer spiral arms.  
Many star-like knots lined up along straight arm.  
Three-armed spiral.

Herc. cluster. (3)  
Paint material from arm to and around comp. Opposite arm faint,  
sweeps around East of galaxy.

Arm leads toward, but not up to companion.  
Broad, diffuse extension of arm leads to companion.  
3C53 is 2°E; M0-013 is 1° E, 26°N. Very faint extension to  
companion.

Apparent gap between arm and companion. (3)  
Material of arm seems to flow "around" comp. Similar to 26(3)(29)  
N at left of picture. Comp. is NGC 770. Faint material toward  
each of two dwarf comps.

Small separation between two knots in arm.  
End of one arm heavy; absorption break in same arm near nucleus.  
Companion resembles M 51 companion. (19)

Arm opposite comp. extremely long.  
Some hazy material at juncture of two arms; high surface brightness,  
S shape inside comp. (35)

Arcs of high S. B. around nucleus of companion.  
Faint plumes and extensions from companion. (3)(25)(33)(47)  
Double arm leading to companion.

TABLE 1 (cont'd)

No.	$\alpha$ (1970)	$\delta$ (1970)	Designation	Plate No.	Exp. (min)	Emul.	Fil- ter	See- ing	Mag	Source	Remarks
87	1 <sup>h</sup> 39.4 <sup>m</sup>	+22 37 <sup>s</sup>	NGC 3808	PH-4368	25	103a-D	-	2	6X	VV300	Position of larger member. Arm appears wrapped around cylindrical comp.
88	1 17.5	+12 19	NGC 2648	4374	25	103a-D	-	3	10X	Wilson 20	Incipient spiral in arm.
89	8 41.0	+14 24	NGC 5930+29	4667	25	103a-J	-	4	4X	Wilson 20	Position of larger spiral. Absorption lanes in comp. Diffuse arm extends beyond comp.
90	15 25.4	+41 46	NGC 5930+54	4486	25	103a-J	-	3	8X	Holmberg 710	N opposite tab mark on picture. Absorption lanes around comp.(34)
91	15 33.2	+15 18	NGC 5933+54	4247	30	103a-D	-	2-3	8X	VV244	N at top of picture. Broad pec. arm to comp., then absorption; faint extension from comp.
92	23 17.3	+ 0 5	NGC 7603	4681	40	103a-E	GG 11	3	8X		Very faint connection shows better in red.
93	22 27.0	-25 0	NGC 7285+84	3993	45	103a-D	-	3	4X	VV74	Long faint plume bifurcates from arm, E comp. in other arm.
94	10 21.8	+20 3	NGC 3226+27	4126	30	103a-D	-	3	2X	VV209	Suggested rotation of axis of spiral.
95	14 33.8	+26 39	IC 4461	4499	25	103a-J	-	2	8X	VV303	Comp. on edge of large, very faint lobe extending opposite galaxy. Light line E-W is plate defect. (25)
96	6 50.8	+86 36		4082	30	103a-D	-	3	6X	VV248	Star-like condensation in spiral. Connection to E galaxy inferred, not seen.
97	12 04.5	+31 14		3891	60	103a-E	GG 11	2-3	8X	VV13	Faint diffuse counter arm, and arm leading to companion.
98	1 30.5	+31 57		4091	30	103a-D	-	3	8X	VV301	High surface brightness S inside spiral, similar to 96.
99	23 13.8	+18 48	NGC 7550	4107	30	103a-D	-	1	2X		Connection not seen, but note difference in arms toward and away from E galaxy. Note also material between West spiral and E galaxy.
100	0 27.0	-11 45	IC 18	4068	30	103a-D	-	2	2X	VV234	Radio source M00-11 is 1.6 E.
101	16 3.1	+14 57		3889	45	103a-D	-	3	4X	VV318	
102	17 18.7	+49 5		3971	50	103a-O	-	3	2X	Zwicky, VV10	VV position wrong. Note loop E side of spiral; diffuse, very faint connection to E galaxy.
103	16 48.8	+45 30	NGC 5216+18	3978	75	103a-O	-	2-3	6X	Zwicky	Incomplete connection, blue knots in southern member.(18)(43)(45)
104	13 31.1	+62 52	NGC 3561	812	30	103a-O	-	1-2	2X	Keenan, VV33	Known as Keenan's system. (32)(44)
105	11 9.6	+28 51		3887	40	103a-D	-	1-2	4X	Zwicky, VV237	Supernova found in disk of spiral. (5)(44)(45)(46)
106	12 14.0	+28 21		3892	60	103a-E	GG 11	3	8X	VV199	Double arm leads to E gal., diffuse material out other side of E galaxy.
107	10 50.6	+30 15		4176	30	103a-D	-	1	6X	VV233	Third arm leads toward E companion.
108	3 1.9	-22 19		4119	30	103a-D	-	2	8X	VV346	
109	15 48.1	+69 31		4256	30	103a-D	-	2	6X	VV291	Arm bent at root.
110	22 52.5	-15 22		4692	24	103a-O	-	4	10X	E galaxy apparently bending arm at root.	
111	14 0.4	+33 58	NGC 5421	4459	35	103a-J	-	3	8X	VV120	
112	23 59.9	+31 17	NGC 7806+05	4001	25	103a-D	lb	4	6X	VV226	
113	0 16.9	+29 55	NGC 70	4277	30	103a-D	-	3	4X	VV166	
114	7 17.9	+85 50	NGC 2276+2300	4418	30	103a-D	GG 11	2	2X		Spiral somewhat pec., may be perturbed. See No. 25.
115	11 41.5	+26 26		4367	30	103a-J	-	2	8X	VV353	Object slightly S of northern gal. is just perceptibly non-stellar.
116	12 42.0	+11 45	NGC 4647+49	367B	30	103a-O	GG 1	2-3	2X	Absorption heavier on spiral side away from E galaxy.	
117	14 8.6	+17 49	IC 982+983	4428	25	103a-D	-	1	4X	Flattening of spiral's nucleus appears to be in different plane than arms.	
118	2 53.6	- 0 17	NGC 1143+44	269	20	103a-O	-	2- $\infty$ 1	10X	Wilson 16, Herzog, VV331	Arms and loops seem attracted to E galaxy.
119	1 17.9	+12 18		4374	25	103a-D	-	3	8X	VV347	Some material seems attracted, some repelled.
120	12 26.3	+13 10	NGC 4438	4425	30	103a-J	-	2	2X	VV188	E galaxy breaking up a spiral. (19)(25)(30)
121	0 57.9	- 4 57		4370	25	103a-D	-	3	8X	Wilson 7	E galaxy warping spiral.
122	16 03.1	+17 46	NGC 6039	4273	30	103a-D	-	2	8X	VV212	Herc. cluster. (3)
123	5 21.1	-11 31	NGC 188+89	4447	30	103a-J	-	2	6X	Page	Faint parallel feature on opposite side from S0 galaxy.
124	17 18.5	+60 38	NGC 6361	4292	45	103a-O	-	1	6X		

TABLE 1 (Cont'd)

No.	$\alpha$ (1970)	$\delta$ (1970)	Designation	Plate No.	Exp. (min)	Fil- ter	Emul.	See- ing	Mag.	Source	Remarks
125	16 37.2 <sup>b</sup>	+41 59 <sup>b</sup>		PH-4503	30	103a-J	-	2	10X	Wilson 37	
126	1 56.6	+ 2 57	NGC 191	4298	30	103a-J	-	2	8X	VV122	(15) Sharp absorption lanes over N side of perturbing galaxy.
127	0 37.4	- 9 10		4660	25	103a-J	-	3	8X		
128	1 16.0	+14 33		4295	30	103a-J	-	1-2	10X	VV205	
129	9 37.7	+32 29		4100	25	103a-D	-	3	10X	VV83	
130	0 1.4	+16 29		4286	30	103a-D	-	3	8X	VV263	
131	2 45.9	-14 54		4388	25	103a-D	-	2	8X	VV336	
132	11 17.9	- 2 56		3225M	25	103a-O	-	2	10X		North inverted. Radio source M11-02 is 3'N. Central member of galaxy group associated with 3C40.
133	1 24.2	- 1 32	NGC 541	4307	60	103a-O	Polaroid	3	8X		
134	12 28.3	+ 8 10	NGC 4472	35B	30	103a-O	-	2	2X	M 49	
135	2 38.5	+38 57	NGC 1023	4345	30	103a-J	-	2	2X	Wilson 15	
136	14 57.8	+64 0	NGC 5820	4630	35	103a-J	-	3	6X	Morgan	
137	9 32.5	+10 14		4363	25	103a-D	-	2	10X	Wilson 21	
138	11 57.2	+25 13	NGC 4015	4396	45	103a-O	-	2	10X	VV216	Absorption leads directly into E galaxy.
139	13 6.0	+26 53		4498	25	103a-J	-	2	10X	Herzog 8	
140	0 49.5	- 7 13	NGC 274-75	4090	30	103a-D	-	3	8X	VV81	
141	7 10.8	+73 32		4083	30	103a-D	-	3	8X	VV123	
142	9 36.2	2 53	NGC 2936+37	4133	30	103a-D	-	3	6X	VV316	
143	7 45.0	+39 11	NGC 2444+45	4084	30	103a-D	-	3	6X	VV117	
144	0 4.4	-13 34	NGC 7828+29	4088	30	103a-D	-	3	8X	VV212	
145	2 21.0	+41 4		163S	30	103a-O	WG2	4	8X	Minkowski, Gates, Reeves	
146	0 5.0	- 6 54		3114M	20	103a-O	-	3-4	10X	Dewhurst Kowal	
147	3 9.6	+ 1 12		4664	25	103a-J	-	4	10X	Mayall, VV82	Known as Mayall's object. (23)
148	11 2.2	+41 0		4353	30	103a-D	-	2	10X	Mayall	
149	12 38.0	+16 46		4478	25	103a-J	-	3	10X	Herzog 42	
150	23 18.0	+ 9 20	NGC 7609	4023Z	25	103a-O	-	2-	10X	VV20	Radio source M23-09 is 2° east.
151	11 23.9	+54 33		4449	25	103a-J	-	3	10X	VV14	
152	12 29.3	+12 34	NGC 4486	363B	10	103a-O	GG 1	2-3	8X	M 87	Short exp. to show jet. Virgo A radio source (3)(9)(32)(33)(34)
153	13 23.6	-42 51	NGC 5128	PS-3272	10	103a-J	Wr. #4	1-2	6X*	Cen A radio source (3)(7)(11)(38)(41) and refs. in Searle.	
154	3 21.5	-37 20	NGC 1316	PH-4684	15	103a-O	-	2	4X	For A radio source. Short exp. to show absorption in center. (20)(38)(41)	
155	11 21.9	+54 1	NGC 3656	4159	30	103a-D	-	2	8X	VV22	
156	10 40.3	+77 37	NGC 520	4380	35	103a-J	-	3	8X	Baade, VV231	Very faint oval loop in NE-SW direction.
157	1 23.0	+ 3 38	NGC 523	1096B	30	103a-D	GG 1	4	2X	Reaves	Note segment in NE direction. (3)
158	1 23.5	+33 52	NGC 4747	4389	25	103a-D	-	2	6X		
159	12 20.3	+25 57	NGC 4194	4535	20	103a-J	-	3	2X		
160	12 12.6	+54 42	NGC 4194	4455	25	103a-J	-	3	6X		
161	11 40.6	+ 0 31		4423	38	103a-J	-	2	8X	Wilson 27	
162	10 49.7	+28 9	NGC 3414	4532	15	103a-J	-	3	6X		
163	12 43.8	+27 17	NGC 4670	289M	20	103a-O	-	1-2	10X		
164	1 14.4	+ 5 2	NGC 455	4296	30	103a-D	-	2	6X		
165	7 34.9	+17 57		4347	30	103a-J	-	2	6X	Wilson 17	
166	1 55.8	+33 4	NGC 750+51	4297	30	103a-D	-	1-2	6X	VV189	
167	8 47.7	+19 12	NGC 2672	4419	15	103a-D	-	3	8X		
168	0 41.1	+40 42	NGC 221	PS-8273	50	103a-J	Wr. #4	4-3	8X*	M 32	Corp. galaxy very condensed, has curved plume.
169	22 13.3	+13 42	NGC 7236+37	PH-4285	30	103a-D	-	2	6X		Faint diffuse plume curved away from M 31 disk.
170	23 15.8	+18 32	NGC 7578	4000	25	103a-D	ib	4	6X	VV181	Faint diffuse plumes coming away from two galaxies. 3C422.
171	14 38.5	+ 3 36		4484	25	103a-J	-	3	8X		M14+010 is 2° west.
172	16 4.2	+17 43		3984	45	103a-D	-	2	6X	VV194	

TABLE 1 (Cont'd)

No.	$\alpha$ (1970)	$\delta$ (1970)	Designation	Plate No.	Exp. (min)	Emul.	Fil- ter	See- ing	Mag	Source	Remarks
173	14 50.0	+9 29	PH-4429	25	103a-D	-	1	6X	VV 296		
174	9 56.7	+28 59	3886	60	103a-E	-	2	4X	Zwicky?		
175	12 31.4	+11 34	IC 3481+83	54Z	40	103a-O	-	1	2X	Zwicky, VV 43	Can see connection only 2/3 way to SE galaxy. (18)(44)
176	13 2.2	-11 20	NGC 4933	4497	25	103a-J	-	2-3	6X		Companion galaxy very condensed.
177	14 54.5	+24 43		4486	25	103a-J	-	3	10X		Very small plume cones off comp. galaxy opposite larger.
178	14 22.9	+34 59		4252	30	103a-D	-	3	6X	VV 77	Ring off center, broad ejected plume from condensation in ring.
179	3 0.3	- 4 49		259M	25	103a-O	-	2	10X		Condensed offset center.
180	4 51.9	- 4 50		279M	20	103a-O	-	1-2	8X		South arm kinks back, thin filament connects nuclei.
181	10 24.2	+80 00	NGC 3210+12	4879	25	103a-D	-	3	6X	VV 319	Long faint filament extends westward from south arm.
182	23 26.3	+ 8 37		4675	120	IIa-O exp.	-	1	6X	VV 343	Long straight, very faint filament like bow wave from comp.
183	13 33.6	+31 33	NGC 1961	4458	30	103a-J	-	3	6X		Three faint patches constitute third arm or filament.
184	5 39.0	+69 25	NGC 6217	4688	30	103a-O	-	3	2X		Two long straight arms or filaments tangent to NE side of galaxy.
185	16 33.8	+78 16	NGC 6217	4676	30	103a-O	-	1-2	4X		Condensed nucleus. Faint outer arms less curved than inner arms.
186	4 32.9	- 8 39	NGC 1614	261M	25	IIa-O	-	<1.1	8X	Radio source M04-012 is 44°W, 31°S; 3C221 is 38°W, 17°S.	Radio source M05-10 is 6°N. Faint filament points to dense nucleus.
187	5 3.5	-10 17		3145M	30	IIa-O	-	1-2	10X		Possible fainter filament toward compact galaxy to NW.
188	16 5.0	+55 37		3977	50	103a-O	-	1-2	4X	Zwicky, VV 29	VV position. Disturbance inside W arm, filament may originate there. (44)
189	12 42.2	+16 33	NGC 4651	165Z	30	103a-O	-	2	2X	VV 56	Radio source near tail apparently not associated. (44)
190	2 48.6	+12 46		4875	35	103a-J	-	3	4X	VV 221	Faint spikes to originate from stellar image; no spectra available.
191	11 5.7	+18 36		4436	30	103a-J	-	2	8X	VV 239	Acute bend in link between galaxies; plumes from stellar-like images.
192	10 35.4	+18 17	NGC 3303	4433	40	103a-D	CG 11	1-2	6X	Wilson 25	Diffuse faint arms off both sides, spike comes from stellar companion.
193	13 19.2	+34 17	IC 883	4483	25	103a-J	-	2-3	8X	Herzog 24	Faint straight outer spikes, hard knots in main body. (23)
194	11 56.2	+36 36		4496	30	103a-O	-	3	8X	VV 126	Outer material connected by thin filament to very hard nucleus.
195	8 52.0	+35 15		4892	30	103a-J	-	2	8X	VV 243	Absorption edge on connection to nucleus.
196	13 13.0	+26 18		4536	25	103a-J	-	3	10X	Herzog 21	Nucleus out of plane of ring. Attachment to companion.
197	11 29.4	+20 38	IC 701	4183	30	103a-D	-	3	8X	VV 3	Straight filament off one end of bar, kink at end of filament.
198	10 58.2	+17 49		4594	30	103a-J	-	2-3	8X	VV 267	Spikes points toward small nucleus; no spectra available.
199	14 15.8	+36 42	NGC 5644+45	384Z	30	103a-O	-	2	10X	VV 210	Spirals appear disturbed.
200	2 52.0	+12 53	NGC 1134	4665	20	103a-J	-	4	10X	VV 38	Splash appearance on W side of gal. points to low S. B. comp. 7'S.
201	0 22.1	- 0 40		4073	30	103a-D	-	3-4	10X	Holmberg 105,	VV connection between galaxies not seen here or on Survey prints.
202	8 58.4	+35 51	NGC 2719	4461	30	103a-J	-	2	6X	Page	Faint tail from smaller galaxy.
203	11 30.1	+28 40	NGC 3712	4493	25	103a-J	-	3	6X		Faint plumes coming off both ends of bar.
204	13 24.0	+84 39		4540	25	103a-J	-	3	8X	VV 39	
205	10 52.9	+54 28	NGC 3448	4552	30	103a-D	-	2	2X	Baade	Much absorption, resolution into stars.
206	10 50.8	+36 48	NGC 3432	4351	30	103a-D	-	2	2X	VV 11	Resolution into stars. Diameter about 0:3 x 0:5.
207	9 27.6	+76 36		4124	30	103a-D	-	2	10X	VV 58	Narrow chaotic absorption tubes across one end.
208	16 50.0	+47 17		4248	30	103a-D	-	2-3	8X	VV 271	Faint straight absorption lanes lead toward nucleus, become triple.
209	16 4.0	+20 38	NGC 6052	4005	45	103a-D	-	3	10X	VV 86, Mayall?	Barred spiral, sharp nucleus, narrow absorption lanes through center. (3)(25)
210	4 28.9	+64 48	NGC 1569	1793B	30	103a-O	GG 1	3-4	6X	Baade	
211	12 35.8	+38 55		4534	30	103a-J	-	3	10X	VV 42	
212	23 19.0	+17 4	NGC 7625	3892	45	103a-D	-	2	8X	Baade	
213	4 4.6	+69 45	IC 356	1607B	30	103a-D	GG 11	3-4	4X	VV 280	
214	11 31.0	+53 14	NGC 3718	1909B	20	103a-D	GG 13	3-4	2X	Baade	
215	9 12.2	+40 14	NGC 2782	1608B	25	103a-D	GG 1	3-4	4X	Baade	
216	23 27.2	+ 3 22	NGC 7678+82	4013	45	103a-D	-	4	2X	VV 329	Patches north of disturbed spiral, emission strong. (25)

TABLE 1 (Cont'd)

No.	$\alpha$ (1970)	$\delta$ (1970)	Designation	Plate No.	Exp. (min)	Fil.	See-	Remarks	
						ter	ing	Mag	Source
217	10 36.9	+53 40	NGC 3310	PH-4364	40	103a-O	-	1-2	2X
218	15 52.3	+18 42		3290	30	103a-D	-	1	10X
219	3 38.4	- 2 13		3142M	25	103a-O	-	1-2	10X
220	15 33.6	+23 35	IC 4583	4502	35	103a-J	-	2	8X
221	9 35.0	-11 11		4167	30	103a-D	-	2-3	8X
222	23 38.3	-12 27	NGC 7727	4002	25	103a-D lb	-	3	2X
223	23 16.4	- 4 49	NGC 7585	4673	30	103a-J	-	2	8X
224	11 49.5	+55 15	NGC 3921	4424	25	103a-D	-	2	4X
225	8 52.1	+78 21	NGC 2655	718B	20	103a-D	GG 1	2	2X
226	22 19.1	-24 50	NGC 7252	4677	40	IIa-O blkd	-	2	2X
227	1 18.5	+ 3 16	NGC 474	4386	35	103a-J	-	3	2X
228	1 47.3	+10 23	IC 162	4069	30	103a-D	-	2	8X
229	1 21.9	+33 6	NGC 5074-08	4314	30	103a-D	-	1	2X
230	0 44.8	-13 37		4369	25	103a-D	-	3	6X
231	0 42.0	- 4 17	IC 1575	4342	30	103a-D	-	1	4X
232	9 32.2	+10 16		4363	25	103a-D	-	2	10X
233	10 30.4	+54 32		54 Bm 20	103a-O	-	1-2	8X	Haro No. 2
234	11 34.2	+54 41	NGC 3738	1138B	30	103a-D	GG 1	4-5	6X
235	0 7.2	+15 39	NGC 14	4072	30	103a-D	-	3	6X
236	1 6.2	-17 38	IC 1623	4357	24	103a-O	-	3	4X
237	9 26.1	+12 25		4462	30	103a-J	-	2	8X
238	(13 13.4)	(+62 18)		4457	30	103a-J	-	3	8X
239	13 40.5	+55 49	NGC 5278+79	383Z	30	103a-O	-	2	VV 19, 383Z
240	13 38.4	+ 0 59	NGC 5257+58	374Z	40	103a-O	-	2	5X
241	14 36.4	+30 35	NGC 4676	4500	25	103a-J	-	2	VV 55
242	12 44.7	+30 54		3790	30	103a-D	GG 13	3	4X
243	8 36.6	+25 52	NGC 2623	1137B	30	103a-O	GG 1	4-5	8X
244	12 0.3	-18 42	NGC 4039+39	422Z	20	103a-O	-	1-2	1X
245	9 44.3	-14 11	NGC 2939+93	4395	25	103a-D	-	2	2X
246	0 5.4	+ 8 12		4384	25	103a-D	-	3	10X
247	8 21.9	+21 26	IC 2338+39	4476	30	103a-J	-	3	8X
248	11 44.9	- 3 26		3893	60	103a-E	GG 11	3	4X
249	23 58.9	+22 50		3396	30	103a-O	-	1-3	VV 35, Wild
250	7 33.8	+35 26		3121M	20	103a-O	-	3-4	10X
251	0 52.1	-14 2		4356	25	103a-D	-	3	10X
252	9 43.4	-19 33		4381	25	103a-D	-	4	8X
253	9 41.9	- 5 9		4135	30	103a-D	-	3	VV 52
254	15 19.9	- 7 15		4501	30	103a-J	-	2	2X
255	9 51.5	+ 8 1		4174	30	103a-D	-	1-2	8X
256	0 17.2	-10 32		3115M	20	103a-O	-	3-4	8X
257	8 50.1	- 2 15		4098	30	103a-D	-	3	8X
258	2 37.4	+18 15		4070	30	103a-D	-	2	VV 41
259	5 0.2	- 4 18	NGC 1741	278M	20	103a-O	-	1-2	6X
260	12 12.1	+16 18		4333	25	103a-J	-	3	Minkowski
261	14 47.8	-10 2		4338	25	103a-J	-	3	8X
262	23 55.3	+16 39		4116	30	103a-D	-	2	Herzog 17
263	10 23.4	+17 19	NGC 3239	4157	30	103a-D	-	1-3	VV 140
									VV 255
									VV 95

TABLE 1 (Cont'd)

No.	$\alpha$ (1970)	$\delta$ (1970)	Designation	Plate No.	Exp. (min)	Emul.	Fil- ter	See- ing	Mag	Source	Remarks
264	10 2.0 12 52.5	+40 ° +36 15	52 NGC 3104 IC 3862	PH-3745S	25	103a-O 4439	GG 13 25	6X 10X	VV 119 VV 266	Faint diffuse outer material. Resolution into knots or knots. Semi-stellar nucleus, faint oval ring outside.	
265	12 57.5 10 35.0	+35 2 +31 43	15 NGC 4861	3746S	25	103a-D 4435	GG 13 35	1-2 2	8X 2X	Resolution into knots, bright knot at S end. (19) Resolution of stars. Note linear loop of emission regions.	
266	8 15.9 12 29.3	+70 50 +41 48	50 NGC 4490+85 NGC 3395+96	4435	30	103a-O 4186	GG 13 30	1-2 3	8X 2X	Semi-stellar nucleus, faint oval ring outside. Resolution of stars, emission loops and dust lanes only.	
267	10 48.2 14 1.8	+33 9 - 5 55	50 NGC 5426+27	4438	30	103a-D 4138	GG 13 30	3	6X VV 246	Resolution of knots, emission loops and dust lanes only. Note arc form of emission knots. (32)(35)	
268	16 4.0 2 19.6	+17 52 +39 14	54 NGC 6054	4251	30	103a-D 4084	GG 13 45	3	4X 10X	Arms linked. Note bifurcation in arm of N spiral. Arms join at dense knot or nucleus. Hero. cluster. (3)(19)	
269	12 26.1 22 18.1	+19 27 +29 14	52 IC 1801+NGC 935 NGC 4809+10	4460	25	103a-D 4071	GG 13 30	2	6X 10X	Position of large spiral. Bright long well defined arms, but smooth, not patchy.	
270	11 36.2 12 40.8	- 2 55 +32 42	55 NGC 4631	4166	30	103a-D 4071	GG 13 30	2	6X VV 323	Perturbation of arm by small galaxy to east.	
271	14 33.5 275	+ 5 29 9 24.4	52 IC 1801+NGC 935 NGC 4809+10	4376	25	103a-D 3812	GG 13 30	3	8X VV 293	Both intersecting edges seem dimmed. Resolution of knots.	
272	12 54.3 276	+ 2 42 26.1	52 NGC 7253	4009	30	103a-D lb IIa-O	GG 13 30	2	6X VV 238	Diffuse material between galaxies, many internal absorption lanes.	
273	3 12.7 1 36.2	- 2 55 +48 4	55 NGC 1253	4079	30	103a-D 4169	GG 13 30	3	2X VV 242	Diffuse material between galaxies, many internal absorption lanes.	
274	1 36.2 12 40.8	+48 4 +32 42	55 NGC 3769	4169	30	103a-D P-8270	GG 13 51	2	6X*	DDO 31	
275	0 35.3 282	+23 50 15.7 +42 5	50 NGC 169	4259	25	103a-D 4099	GG 13 30	2	6X VV 50	Knots resolved with 48-inch. Diffuse counter tail on companion. (3)(25)(29)(33)(40)	
276	23 34.7 9 22.1	+ 2 0 +49 23	50 NGC 2798+99	4099	30	103a-D lb 4012	GG 13 30	3	6X VV 51	Companion appears to rain into nucleus of spiral.	
277	14 18.8 9 1.0	+ 4 6 +26 3	56 NGC 5666+60+69	4448	30	103a-J 4088	GG 11 45	4	4X VV 15	Arc of barely resolved knots curves into nucleus of larger galaxy. Some very small knots in connecting streamer. (32)	
278	13 33.5 1 54.7	+13 58 -19 44	56 NGC 5221+22+26	4123	30	103a-D 4127	GG 11 35	2	4X VV 40	Narrow tail leads away from northern nucleus. Connected not visible.	
279	2 2.3 10 41.2	+14 35 +13 37	56 IC 196+95	4182	30	103a-D 4239	GG 11 30	3	2X VV 40	Slanted parallel streamers off each edge of main galaxy. Streamers in both directions from edge of spiral.	
280	9 53.1	- 6 43	57 IC 575	4463	17	103a-J 4136	GG 11 30	1	2X VV 112	Very faint diffuse streamers. VV position incorrect.	
281	16 58.1	+58 58	58 NGC 3786+88	4291	30	103a-D 4099	GG 11 30	3	8X VV 111	Main body has cylindrical appearance. Position of central object. Edge-on Sa, some indication of absorption streaming off edges.	
282	9 15.7 283	+42 5	58 NGC 2798+99	4385	25	103a-D 4099	GG 11 30	3	6X VV 50	Position of larger galaxy. Companion NW. Diffuse arc SE of brighter galaxy.	
283	23 34.7	+ 2 0	58 NGC 7114+15	4012	30	103a-D lb 4012	GG 11 30	4	4X VV 51	Peculiar filaments. (35)	
284	9 22.1	+49 23	58 NGC 2854+56	4448	30	103a-J 4088	GG 11 45	2	4X VV 15	SW gal. is IC 1505. Polarized bridge. (6)(32)(44)	
285	14 44.1	+ 4 6	58 NGC 5566+60+69	4088	30	103a-D 4123	GG 11 30	2	2X VV 40	Long st. filament almost to attachment with arm of spiral.	
286	9 1.0	+26 3	58 NGC 5221+22+26	4127	35	103a-J 4127	GG 11 35	1-2	2X VV 315	Position of larger spiral. Companion on arm has long tail extending westward.	
287	13 33.5 289	+13 58 -19 44	58 NGC 3981	4182	30	103a-D 4239	GG 11 30	3	2X VV 8	Absorption, knots. Note apparent re-entrant spiral arm on southern galaxy.	
288	2 2.3 290	+14 35 +13 37	58 IC 196+95	4463	17	103a-J 4136	GG 11 30	2	6X VV 309	Bright internal knots.	
289	10 41.2	+13 37	58 IC 575	4136	30	103a-D 4291	GG 11 30	3	8X VV 112	Position between pair. Note elongated feature pointing toward nucleus of larger spiral.	
290	9 53.1	- 6 43	58 NGC 3786+88	4291	30	103a-D 4099	GG 11 30	1	6X VV 228	Position of larger spiral.	
291	11 27.3	+58 43	58 NGC 7469	4376	26	103a-O 4165	GG 11 30	2	2X VV 118	Peculiar filaments. (35)	
292	1 36.2 293	+ 8 42	58 NGC 4016+17	4354	30	103a-D 4165	GG 11 30	3	6X VV 106	SW gal. is IC 1505. Polarized bridge. (6)(32)(44)	
293	16 58.1	+58 58	58 NGC 3786+88	4291	30	103a-D 4099	GG 11 30	2	8X VV 229	Long st. filament almost to attachment with arm of spiral.	
294	11 38.1	+32 5	58 IC 1505	373Z	40	103a-O 3980	GG 11 75	2	2X VV 34	Position between pair.	
295	23 39.9	- 3 44	58 NGC 3690+IC 694	4354	30	103a-O 4359	GG 11 30	2	6X VV 34	Zwicky	
296	11 27.1	+58 42	58 NGC 5754+55	4359	30	103a-D 4088	GG 11 75	3	2X VV 34	Wirtanen 17	
297	14 44.1	+38 52	58 NGC 4016+17	4359	30	103a-J 4088	GG 11 75	3	2X VV 34	Wilson 31	
298	23 1.7	+ 8 42	58 NGC 7469	3976	26	103a-O 4165	GG 11 30	3	6X VV 106	Segment breaking from arm of S gal., weak filaments reach to N gal., which has figure 8 loops.	
299	11 27.3	+58 43	58 NGC 4016+17	4354	30	103a-D 4165	GG 11 30	2	6X VV 106	VV position. Resolution; diffuse, hooked countertail.	
300	9 25.4	+68 32	58 NGC 3786+88	4354	30	103a-D 4165	GG 11 30	3	4X VV 173+174	Position between pair. Possibly not interacting.	
301	11 8.3	+24 27	58 IC 563+64	4434	25	103a-D 4485	GG 11 25	2	8X VV 340		
302	14 55.7	+24 44	58 IC 1241+42	4435	25	103a-J 4491	GG 11 25	3	8X VV 340		
303	9 44.7	+ 3 12	58 NGC 4016+17	4435	25	103a-J 4509	GG 11 30	3	6X VV 334		
304	3 9.8	- 9 2	58 NGC 4016+17	4435	25	103a-D 4453	GG 11 25	2	2X VV 334		
305	11 57.0	+27 44	58 NGC 2872+74	4077	30	103e-D 4492	GG 11 25	3	4X VV 106		
306	1 31.0	+ 4 27	58 NGC 2872+74	4077	30	103e-D 4492	GG 11 25	3	6X VV 106		
307	9 24.2	+11 34	58 NGC 2872+74	4077	30	103e-D 4492	GG 11 25	3	4X VV 106		

TABLE 1 (Cont'd)

No.	$\alpha$ (1970)	$\delta$ (1970)	Designation	Plate No.	Exp. (min.)	Emul.	Fil- ter	See- ing Mag.	Source	Remarks
308	1 24.2	- 1 31	NGC 545+47	PH-4307	60	103a-O	Polaroid	3	8X	Closely elliptical. Position of central galaxy (NGC 541).
309	2 26.9	-10 58	NGC 942+43	4118	30	103a-D	-	3	6X	Peculiar absorption ring, possibly broken.
310	17 26.9	+58 33	IC 1259	3985	45	103a-D	-	3	10X	Very close E galaxies. Picture is 10X of following (No. 311) area.
311	17 26.9	+58 33	IC 1259	3985	45	103a-D	-	3	4X	Same as 310, but shows surrounding field and group. Picture 1 $\frac{1}{8}$ 4X of preceding (No. 310) area.
312	16 49.0	+46 45	NGC 3994+95	4283	30	103a-D	-	2	8X	Diffuse connection between central members of group.
313	11 56.1	+32 27	NGC 3994+95	4170	30	103a-D	-	3	2X	Linear strings of knots like deformed spiral arms. Strong [O II].
314	22 56.3	- 3 57	NGC 2832	3974	50	103a-O	-	3	2X	Faint filament leads SE to faint dwarf. Pos. of W spiral.
315	9 18.0	+33 54	NGC 3190	4432	30	103a-D	GG 11	2	8X	Companion E is quite compact.
316	10 16.5	+21 59	PS-8268	50	103a-J	Wr. #4	2	8X*	Edge-on spiral shows signs of interaction.	
317	11 18.6	+13 10	NGC 3627+23+28	PS-8269	50	103a-J	Wr. #4	2	2X*	See also 16. Both galaxies on east show signs of interaction.
318	2 7.9	-10 16	NGC 833+38+39	PH-4372	30	103a-J	-	4	2X	Position of NGC 833. Faint, diffuse streamers, peculiar galaxies.
319	22 34.5	+33 47	NGC 7317 thru 19	4657	20	103a-J	-	3	2X	Position of NGC 7317. Stefan's Quintet. (13)(18)
320	11 36.3	+22 11	4444	25	103a-D	-	2	4X	Position of close triplet. Large companion NW.	
321	9 37.4	- 4 42	37255	25	103a-O	GG 13	2	8X	VV position. Sharp absorption lane in connection to southern most galaxy. (18)	
322	11 31.1	+53 7	1909B	20	103a-O	GG 13	3	8X	Near NGC 3718. See No. 214. (18)	
323	23 52.6	+ 0 13	NGC 7783	4661	20	103a-J	-	4	6X	VV 150
324	16 0.8	+16 2	PS-8521	50	103a-J	Wr. #4	2	10X*	VV 208	
325	22 1.7	-21 12	PH-4302	30	103a-O	-	3	10X	VV 159	
326	13 36.1	+ 6 35	4537	25	103a-J	-	3	2X	Position of integral sign spiral. Five spirals in approx. chain. No. 33 gives larger scale picture.	
327	5 20.2	+ 6 39	NGC 1875	4666	20	103a-J	-	4	8X	Three distorted galaxies in general line toward east.
328	14 46.5	+19 11	4487	35	103a-J	-	3	4X	Six galaxies more or less in line; center one has semi-stellar component.	
329	11 30.4	+70 58	4450	35	103a-J	-	2	8X	Diffuse elongation of E's along line joining them.	
330	16 48.3	+53 27	4654	20	103a-E	GG 11	2	2X	Position of integral sign spiral. Five spirals in approx. chain.	
331	1 5.7	+32 15	NGC 375 thru 388	4668	25	103a-E	GG 11	2	1X	Makarian
332	3 7.2	-23 10	IC 1892	4373	25	103a-J	-	3	IX	Different types of galaxies in chain.
333	2 37.6	+10 43	NGC 1024	4318	45	103a-O	-	1	4X	Thin circular arms, star in SE superposed on wisp.
334	13 29.1	+31 47	4440	25	103a-D	-	2	6X	Second "star" south not quite stellar.	
335	11 1.8	+ 4 56	NGC 3509	4180	30	103a-D	-	3	8X	Large luminous system. (19)
336	8 53.8	+58 55	NGC 2685	6638	30	103a-O	-	2	3X	E is to right of N, W to left. (3)(14)(26)
337	9 53.6	+69 51	NGC 3034	5Bm	30	103a-O	WG 2	3	2X	Internal explosion. (3)(8)(20)(25)(29)(37)
338	10 9.5	- 7 46	3233M	20	103a-O	-	1-2	10X	M 82	

TABLE 2  
ATLAS REDSHIFTS

$\alpha$ (1970) h m	$\delta$	No.	Designation	Corrected Redshift km/sec	$\alpha$ (1970) h m	$\delta$	No.	Designation	Corrected Redshift km/sec
0 1.4	+16 29	130			2 2.3	+14 35	290	IC 195+96	
0 4.2	-13 36	51	NGC 7828+29		2 6.4	+41 20	74		
0 4.4	-13 34	144	NGC 7828+29		2 7.9	-10 16	318	NGC 833+35+38+39	
0 5.0	-6 54	146			2 16.6	+ 5	30	10	
0 5.4	+ 8 12	246	NGC 144		2 19.6	+39 14	273		
0 7.2	+15 39	235	NGC 144		2 21.0	+41 4	145		
0 16.9	+29 55	113	NGC 70		2 23.0	- 4	47	54	
0 17.2	-10 32	256			2 26.1	+19 27	276	IC 1801, NGC 935	
0 20.3	+22 13	65			2 26.9	-10 58	309	NGC 942+43	
0 20.7	- 1 34	35			2 37.4	+18 15	254		
0 22.1	- 0 40	201	IC 18		2 37.6	+10 43	333	NGC 1024	
0 27.0	-11 45	100			2 38.5	+38 57	135	NGC 1023	
0 30.2	- 5 19	19	NGC 145		2 41.1	- 0 8	37	NGC 1068	
0 35.3	+23 50	282	NGC 169		2 45.2	-30 24	77	NGC 1097	
0 37.4	- 9 10	127	NGC 191		2 45.9	-14 54	131		
0 41.1	+40 42	168	NGC 221	+17, +38 (1)	2 48.6	+12 46	190		
0 42.0	- 4 17	231			2 52.0	+12 53	200	NGC 1134	
0 44.8	-13 37	230			2 53.6	- 0 17	118	NGC 1133+44	
0 49.5	- 7 13	140	NGC 274+75	+1854, +1971 (4)	3 0.3	- 4 49	179		
0 52.1	-14 2	251			3 1.9	-22 19	108		
0 57.9	- 4 57	121			3 7.2	-23 10	332	IC 1892	
0 59.0	- 9 19	59	NGC 341		3 8.4	-20 41	41	NGC 1232	
1 5.7	+32 15	331	NGC 333		3 9.6	+ 1 12	147		
1 6.2	-17 38	236	IC 1623		3 9.8	- 9 2	304	NGC 1241+42	
1 7.9	+14 11	11			3 12.7	- 2 55	279	NGC 1253	
1 14.4	+ 5 2	164	NGC 465		3 21.5	-37 20	154	NGC 1316	
1 16.0	+14 33	128			3 28.5	-22 22	39	NGC 1347	
1 17.5	+12 19	88			3 38.4	- 2 13	219		
1 17.9	+12 18	119			4 4.6	+69 45	213	IC 356	
1 18.5	+12 11	48	NGC 474	+2402 (1)	4 18.4	+ 2 1	20		
1 18.5	+ 3 16	227			4 32.9	- 8 39	186	NGC 1614	
1 19.8	- 0 42	67			4 35.2	- 2 21	61		
1 20.9	- 1 1	8	NGC 497		4 28.9	+64 48	210	NGC 1569	
1 21.8	+30 37	70			4 51.9	- 4 50	180		
1 21.9	+33 6	229	NGC 507+08	+5121 (1), +2320 (4)	5 0.2	- 4 18	259	NGC 1741	
1 23.0	+ 3 38	157	NGC 520		5 3.5	-10 17	187		
1 23.5	+33 52	158	NGC 523		5 18.0	+ 3 42	52		
1 24.2	- 1 32	133	NGC 541		5 20.2	+ 6 39	327	NGC 1875	
1 24.2	- 1 31	308	NGC 545+47		5 21.1	-11 31	123	NGC 1883-89	
1 30.5	+31 57	98			5 39.0	+69 25	184	NGC 1961	
1 31.0	+ 4 27	306			6 50.8	+86 36	96		
1 47.0	-12 31	4			7 10.8	+73 32	141		
1 47.3	+10 23	228	IC 162		7 17.9	+85 50	25	NGC 2276	
1 49.4	+21 45	31			7 17.9	+85 50	114	NGC 2276+2300	
1 49.8	- 4 12	75	NGC 702		7 33.8	+35 26	250		
1 55.6	+17 4	56							
1 55.8	+33 4	166	NGC 750+51	+5291 (1), +5295 (1)					
1 56.6	+ 2 57	126							
1 57.7	+18 52	78	NGC 772	+2553 (1)					

TABLE 2 (Cont'd)

$\alpha$ (1970) $\delta$				No.	Designation	Corrected Redshift km/sec	$\alpha$ (1970) $\delta$	No.	Designation	Corrected Redshift km/sec		
h	m	.	s				h	m	.	s		
7	34.9	+17	57	165			10	21.8	+20	3	94	NGC 3226+27
7	41.3	+73	52	17			10	22.3	+16	53	43	
7	45.0	+39	11	143	NGC 2444+45	+3951 (4)	10	23.4	+17	19	263	NGC 3239
8	9.4	+25	18	82	NGC 2535+36	+4153 (1), +3983 (4), +4046 (4)	10	24.2	-2	3	44	IC 609
8	11.1	+46	5	6	NGC 2537	+415, +299 (1), +422 (4)	10	20.4	+80	0	181	NGC 3210+12
8	11.6	+73	42	9	NGC 2523	+3604 (1)	10	33.6	+54	32	233	NGC 3290
8	15.9	+70	50	268			10	35.0	+31	43	267	
8	21.9	+21	26	247	IC 2388+39		10	35.4	+18	17	192	NGC 3303
8	30.2	+19	19	58			10	36.9	+53	40	217	NGC 3310
8	33.4	+28	41	12	NGC 2608	+2041 (1)	10	40.3	+77	37	156	
8	36.6	+25	52	243	NGC 2623	+5342 (1)	10	41.2	+13	37	291	
8	41.0	+14	24	89	NGC 2648	+2382 (4)	10	48.2	+33	9	270	NGC 3395+96
8	45.0	+74	14	80	NGC 2623	+4100 (1)	10	49.7	+28	9	162	NGC 3414
8	47.7	+19	12	167	NGC 2672		10	50.6	+30	15	107	
8	49.1	-16	30	7			10	50.8	+36	48	206	NGC 3432
8	50.1	-2	15	257			10	52.8	+57	9	24	NGC 3445
8	52.0	+35	15	195			10	52.9	+54	28	205	NGC 3448
8	52.1	+78	21	225	NGC 2655	+1473 (1)	10	58.2	+17	49	198	
8	53.8	+58	55	336	NGC 2685	+961 (1), +957 (4)	11	1.8	+ 4	56	335	NGC 3509
8	58.4	+35	51	202	NGC 2719	+3073 (4)	11	2.2	+41	0	148	
9	1.0	+26	3	287			11	3.2	+30	15	21	
9	12.2	+40	14	215	NGC 2782	+2502 (1), +2514 (4)	11	5.7	+18	36	191	
9	13.9	+44	27	55			11	8.3	+24	27	301	
9	15.7	+42	5	283	NGC 2798+99	+1699 (1)	11	9.6	+28	51	105	NGC 3561
9	18.0	+33	54	315	NGC 2832	+6895 (1)	11	17.9	- 2	56	132	
9	22.1	+49	23	285	NGC 2854+56		11	18.6	+13	10	317	NGC 3627
9	22.9	+49	28	1	NGC 2857		11	18.6	+13	11	16	NGC 3631
9	24.2	+11	34	307	NGC 2872+74	+2803 (4), +3443 (4)	11	19.4	+53	21	27	
9	24.4	-11	52	275	NGC 2881		11	21.9	+54	1	165	NGC 3656
9	25.4	+68	32	300			11	22.7	+ 3	29	5	NGC 3664
9	26.1	+12	25	237			11	23.9	+54	33	151	
9	27.6	+76	36	207			11	27.1	+58	42	296	NGC 3690+IC 694
9	32.2	+10	16	232			11	27.3	+58	43	299	
9	32.5	+10	14	137			11	29.4	+20	38	197	IC 701
9	35.0	-11	11	221			11	30.1	+28	40	203	NGC 3712
9	36.2	+ 2	53	142	NGC 2936+37		11	30.4	+70	58	329	
9	37.4	+32	27	63	NGC 2944		11	31.0	+53	14	214	NGC 3718
9	37.4	- 4	42	321			11	31.1	+53	7	322	
9	37.7	+32	29	129			11	34.2	+54	41	234	NGC 3738
9	41.9	- 5	9	253			11	36.2	+48	4	280	NGC 3769
9	43.4	-19	33	252			11	36.3	+22	11	320	
9	44.3	-14	11	245	NGC 2992+93		11	44.9	- 3	26	248	
9	44.7	+ 3	12	303			11	49.5	+55	15	224	NGC 3921
9	51.5	+ 8	1	255			11	58.6	+15	29	83	NGC 3786+88
9	53.1	- 6	43	292	IC 575		11	54.7	-19	44	62	NGC 3981
9	53.6	+69	51	337	NGC 3034	+400, +410 (1), +322 (4)	11	54.1	+22	37	87	NGC 3799+3800
9	56.7	+28	59	174			11	54.6	+ 0	31	161	
10	2.0	+40	52	264	NGC 3104		11	41.5	+26	26	115	
10	9.5	- 7	46	338			11	44.9	- 3	26	248	
10	16.5	+21	59	316	NGC 3190	+1220, +1254 (1), +1255 (4)	11	56.1	+32	27	313	NGC 394+95

TABLE 2 (Cont'd)

$\alpha$ (1970) $\delta$				Designation				Corrected Redshift km/sec				$\alpha$ (1970) $\delta$				Designation				Corrected Redshift km/sec				
h	m	o	'	No.				h	m	o	'	No.				h	m	o	'	No.				
11	56.2	+36	36	194				14	2.1	+54	29	26	NGC 5457				14	8.5	+17	46	79	IC 982+983		
11	57.0	+27	44	305	NGC 4016+17			14	8.6	+17	49	117	NGC 5544+45				14	15.8	+36	42	199	+32265, +3275 (4)		
11	57.2	+25	13	138	NGC 4015			14	18.7	+52	0	45					14	18.8	+4	6	286	NGC 5560+66+69	+1436 (1), +1581 (4)	
11	58.0	-19	6	22	NGC 4027	+1701 (4)		14	19.2	+35	19	69	NGC 5579+80				14	22.9	+34	59	178			
12	0.3	-18	42	244	NGC 4038+39	+1469 (1), +1443 (4), +1456 (1), +1427 (4)		14	30.9	+8	12	49	NGC 5665				14	30.9	+5	29	274			
12	4.0	+50	42	18	NGC 4088	+820 (1), +812 (4)		14	33.5	+26	39	95	IC 4461				14	36.4	+30	35	241			
12	4.5	+31	14	97				14	38.5	+3	36	171				14	43.8	+19	36	64	NGC 5754+55			
12	12.1	+16	18	260	NGC 4194	+2684 (1)		14	44.1	+38	52	297				14	44.1	+38	52	297				
12	12.6	+54	42	160				14	45.8	+18	59	47				14	45.8	+19	11	328				
12	14.0	+28	21	106	NGC 1020	NGC 4438	-105 (1), +855 (4)	14	46.5	+19	11					14	47.8	-10	2	261				
12	26.3	+13	10	120	NGC 4472	+918 (1), +846 (4)	+675 (1), +848 (4), +622 (4)	14	50.0	+9	29	173				14	54.5	+24	43	177				
12	28.3	+8	10	134	NGC 4486	+1218, +1187 (4), +1124 (1)	+1218, +1187 (4), +1124 (1)	14	55.7	+24	44	302	NGC 5820	+3444 (1)		14	55.7	+24	44	302	NGC 5829			
12	29.3	+41	48	269	NGC 4486+90	+7011 (1), +31 (4)	+896 (1)	14	57.8	+54	0	136				14	59.7	+23	27	42				
12	29.3	+12	34	152	NGC 3481+83	+896 (1)		15	1.4	+23	27	42				15	19.9	-7	15	254				
12	31.4	+11	34	175	NGC 4569			15	19.9	+54	0	136				15	25.4	+41	46	90	NGC 5930	+2668 (4)		
12	35.3	+13	19	76				15	33.2	+15	38	91	NGC 5953+54	+2188, +2228 (4)		15	33.6	+23	39	220	IC 4553			
12	35.8	+38	55	211				15	45.7	+18	00	72	NGC 5594+96			15	48.1	+69	31	109				
12	38.0	+16	46	149				15	52.3	+18	42	218				16	0.8	+16	2	324				
12	40.0	+26	14	34	NGC 4618	+541 (1)	+611 (1), +646 (4)	+1321, +1379 (1), +1328 (4), +1175 (1), +1200 (4)	16	3.1	+14	57	101	NGC 6039			16	3.1	+17	46	122	NGC 6054		
12	40.1	+41	19	23	NGC 4631	+611 (1), +646 (4)	+1321, +1379 (1), +1328 (4), +1175 (1), +1200 (4)	16	4.0	+17	46	71	NGC 6052			16	4.0	+20	38	209	NGC 6052			
12	40.8	+32	42	281	NGC 4647+49	+1209 (1)	+6515 (A), +6320 (B), (4)	+1209 (1)	16	3.8	+17	46	71			16	5.0	+55	37	188				
12	42.0	+11	45	116	NGC 4651	+6515 (A), +6320 (B), (4)	+6515 (A), +6320 (B), (4)	16	15.1	+47	7	2			16	26.1	+51	36	66	NGC 6217	+1617, +1613 (1)			
12	42.2	+16	33	189	NGC 4670	+6515 (A), +6320 (B), (4)	+6515 (A), +6320 (B), (4)	16	33.8	+17	52	72	IC 1222			16	33.8	+78	16	185				
12	43.8	+27	17	163	NGC 4670	+6515 (A), +6320 (B), (4)	+6515 (A), +6320 (B), (4)	16	43.0	+17	52	72			16	34.2	+46	17	73					
12	44.7	+30	54	242	NGC 4676	+854 (4), +794 (4)	+829 (1), +831 (4)	16	34.2	+41	58	293			16	34.2	+46	17	208					
12	50.3	+25	57	159	NGC 4747	+854 (4), +794 (4)	+829 (1), +831 (4)	16	35.2	+41	58	293			16	35.2	+46	17	208					
12	52.5	+36	15	265	IC 3862			16	33.6	+23	39	330			16	33.6	+23	39	330					
12	54.3	+2	42	277	NGC 4809+10	+829 (1), +831 (4)	+829 (1), +831 (4)	16	48.1	+69	31	109			16	48.1	+69	31	109					
12	57.5	+35	2	266	NGC 4861			16	52.3	+18	42	218			16	52.3	+18	42	218					
13	2.2	-11	20	176	NGC 4933			16	0.8	+16	2	324			16	0.8	+16	2	324					
13	6.0	+26	53	139				16	3.1	+14	57	101			16	3.1	+14	57	101					
13	13.0	+26	18	196				16	4.0	+17	43	172			16	4.0	+20	38	209					
13	13.2	+26	16	60				16	5.0	+55	37	188			16	5.0	+55	37	188					
13	13.4	(+62	18)	238				16	15.1	+47	7	2			16	15.1	+47	7	2					
13	15.2	+14	35	57	IC 883	+6918 (23)	+261 (1), +271 (4)	+6918 (23)	16	26.1	+51	36	66			16	33.8	+78	16	185				
13	19.2	+34	17	193				16	34.2	+46	17	208			16	34.2	+46	17	208					
13	23.6	-42	51	153	NGC 5128			16	35.2	+41	58	293			16	35.2	+41	58	293					
13	24.0	+84	39	204	IC 4241			16	48.3	+53	27	330			16	48.3	+53	27	330					
13	28.1	+37	34	40				16	48.8	+45	30	103			16	48.8	+45	30	103					
13	28.6	+47	21	85	NGC 5194+95	+546 (1), +552 (4)	+546 (1), +552 (4)	+546 (1), +552 (4)	16	49.0	+46	45	312			16	49.0	+46	45	312				
13	29.1	+31	47	334	NGC 5216+18	+6744, +6569 (4)	+6744, +6569 (4)	+6744, +6569 (4)	16	50.0	+47	17	208			16	50.0	+47	17	208				
13	31.1	+62	52	104	NGC 5257+58	+7665, +7708 (4), +3651 (1)	+7665, +7708 (4), +3651 (1)	+7665, +7708 (4), +3651 (1)	16	58.1	+58	58	293			16	58.1	+58	58	293				
13	32.9	+31	35	36	NGC 5221+22+26				16	58.1	+58	58	293			16	58.1	+58	58	293				
13	33.5	+13	58	288				16	59.1	+60	38	124	NGC 6361			16	59.1	+60	38	124				
13	33.6	+31	33	183				16	60.5	+61	39	125			16	60.5	+61	39	125					
13	36.1	+6	35	33				16	61.9	+62	40	126			16	61.9	+62	40	126					
13	38.4	+0	59	240	NGC 5257+58				16	63.3	+64	41	127			16	63.3	+64	41	127				
13	40.5	+55	49	239	NGC 5278+79				16	64.7	+65	42	128			16	64.7	+65	42	128				
13	57.4	+37	35	84	NGC 5394+95	+3651 (1)	+3651 (1)	+3651 (1)	16	66.1	+66	43	129			16	66.1	+66	43	129				
14	0.4	+33	58	111	NGC 5421				16	67.5	+67	44	130			16	67.5	+67	44	130				
14	1.8	-5	55	271				16	68.9	+68	45	131			16	68.9	+68	45	131					

TABLE 2 (Cont'd)

$\alpha$ (1970)	$\delta$	No.	Designation	Corrected Redshift km/sec	$\alpha$ (1970)	$\delta$	No.	Designation	Corrected Redshift km/sec
h m	° '				h m	° '			
17 18.7	+49 5	102			22 58.6	+15 49	13	NGC 7448	+2649 (1)
17 22.5	+62 11	30	IC 1259		23 1.7	+ 8 42	298	NGC 7469	+5015 (4), +4988, +4899 (1)
17 26.9	+58 33	310	IC 1259		23 13.8	+18 48	99	NGC 7550	
17 26.9	+55 33	311	NGC 6412	+1751 (1)	23 15.8	+18 32	170	NGC 7578	
17 30.1	+75 45	38			23 16.4	- 4 49	223	NGC 7586	+3502 (4), +3485, +3538 (1)
18 13.4	+68 18	81	NGC 6621+22	+6490 (4)	23 17.3	+ 0 5	92	NGC 7603	
					23 18.0	+ 9 20	150	NGC 7609	
20 34.3	+60 2	29	NGC 6946	+371 (4), +330, +221 (1)	23 19.0	+17 4	212	NGC 7625	+2050 (1), +2009 (4)
22 1.7	-21 12	325			23 26.3	+ 8 37	182		
22 13.3	+13 42	169	NGC 7286+37	+8038, +8093 (4)	23 27.1	+22 15	28	NGC 7678	+3676 (1), +3680 (4)
22 18.1	+29 14	278	NGC 7253		23 27.2	+ 3 22	216	NGC 7679+82	+8378, +5278 (1), +5330 (4)
22 19.1	-24 50	226	NGC 7252	+4817 (4)	23 32.2	+29 52	46		
22 27.0	-25 0	93	NGC 7284+85		23 34.7	+ 2 0	284	NGC 7714+15	+3001 (1), +2963 (1)
22 34.1	-26 12	14	NGC 7314	+1838 (1)	23 38.3	-12 27	222	NGC 7727	+1953 (4), +1943, +1982 (1)
22 34.5	+33 47	319	NGC 7317-19	(1), (13)	23 39.9	- 3 44	295	IC 1505	
22 34.8	- 3 6	3			23 45.6	+29 19	86	NGC 7752+53	
22 50.0	- 5 43	15	NGC 7393	+3975 (1), +3972 (4)	23 47.0	+ 3 57	68	NGC 7756	+5108, +5085 (4)
22 52.5	-15 22	110			23 52.6	+ 0 13	323	NGC 7783	
22 56.3	- 3 57	314			23 55.3	+16 39	262		
					23 56.4	-14 12	50	IC 1520	
					23 58.9	+22 50	249		
					23 59.9	+31 17	112	NGC 7805+06	

TABLE 3  
RADIO OBSERVATIONS ON OR NEAR ATLAS GALAXIES

Atlas	Designation	Radio Source Name	References	Remarks	References
331	NGC 383	3C31	(1)(5)	Diameter $2.5 \pm 1.5$	(1) Edge, D. O., Shakeshaft, J. R., McAdam, W. B., Baldwin, J. E., and Archer, S., R. A. S. Memoirs, LXVIII, 37.
227	NGC 470+74	M01-03	(1)(0)(1)	Also 3C39?	(2) Wade, C. M., Pub. Nat. Radio Astr. Obs., Vol. 1, No. 6.
133, 308	NGC 541+45+47	3C40, M01-01	(1)(5)(7)(10)	Double source	(3) Heeschen and Wade, A. J., 69, 277 (1969).
78	NGC 772	(3, Table IV)			(4) Clarke, Margaret E., M. N., 127, 405 (1964).
37	NGC 1068	M02-00, 3C71	(1)(3)(7)	Seyfert Galaxy	(5) Bennet, A. S., 1962, Mem. R. A. S. LXVII, 163.
164	NGC 1316	M03-37	(6)(2)(11)	For A	(6) Stanley, G. J. and Slee, O. B., 1950, Astr. J. Sci. Res. A, 3, 234.
283	NGC 2798-9	(3, Table X, XII)	(10)	VV50	(7) Fomalont, E. B., Matthews, T. A., Morris, D., and Wyndham, J. D.
252	NGC 3034	M09-19	(1)(4)(3, Table IV)(7)(8)	M82	(8) Wyndham, J. D., and Read, R. B., 1965, A.J., 70, 120.
337	NGC 3310	3C231	(3, Table X)		(9) Bolton, J. G., 1948, Nature 162, 141.
217	NGC 3448	(3, Table X)			(10) Mills, B. Y., Slee, O. B., and Hill, E. R., 1958, Astr. J. Physics 11, 362.
205	NGC 3627+28	(3, Table IV, XII)		VV308	(11) Mills, B. Y., Slee, O. B., and Hill, E. R., 1960, Astr. J. Physics 13, 676.
16, 317	NGC 4038+39	M11-18	(3, Table IV, XII)	VV245	
244	NGC 4088	(3, Table IV)			
134	NGC 4472	(3, Table IV)			
152	NGC 4486	3C274	(3, Table IV)(1)(7)(8)	$\Delta\alpha = 0^{\circ}8'$	
269	NGC 4490-85	(3, Table XII)		M87, Vir A	
76	NGC 4569	(3, Table IV)		VV30	
281	NGC 4631	(3, Table IV)			
189	NGC 4651	3C275.1	(3, Table IV)(1)		
153	NGC 5128	M13-42	(9)(2)(7)	Prob. not assoc. w spiral	
85	NGC 5194	(3, Table IV, XII)		Cen A	
84	NGC 5394-5	(3, Table XII)		M51, VV1	
26	NGC 5457	(3, Table IV, XII)		VV48, Possible radio source	
286	NGC 5566	M14-04	(10)	M101, VV344	
29	NGC 5946	(3, Table IV)			
169	NGC 7238+37	3C442	(5)(7)(8)	$4' < \text{diam} < 10'$	
298	NGC 7469	(3, Table X)		Confused region	
	See also remarks after Atlas No. 35, 75, 150, 171, 186, 132, 256,			100, 53	

MSH sources designated with first two figures from right ascension plus sign and first two figures from declination.

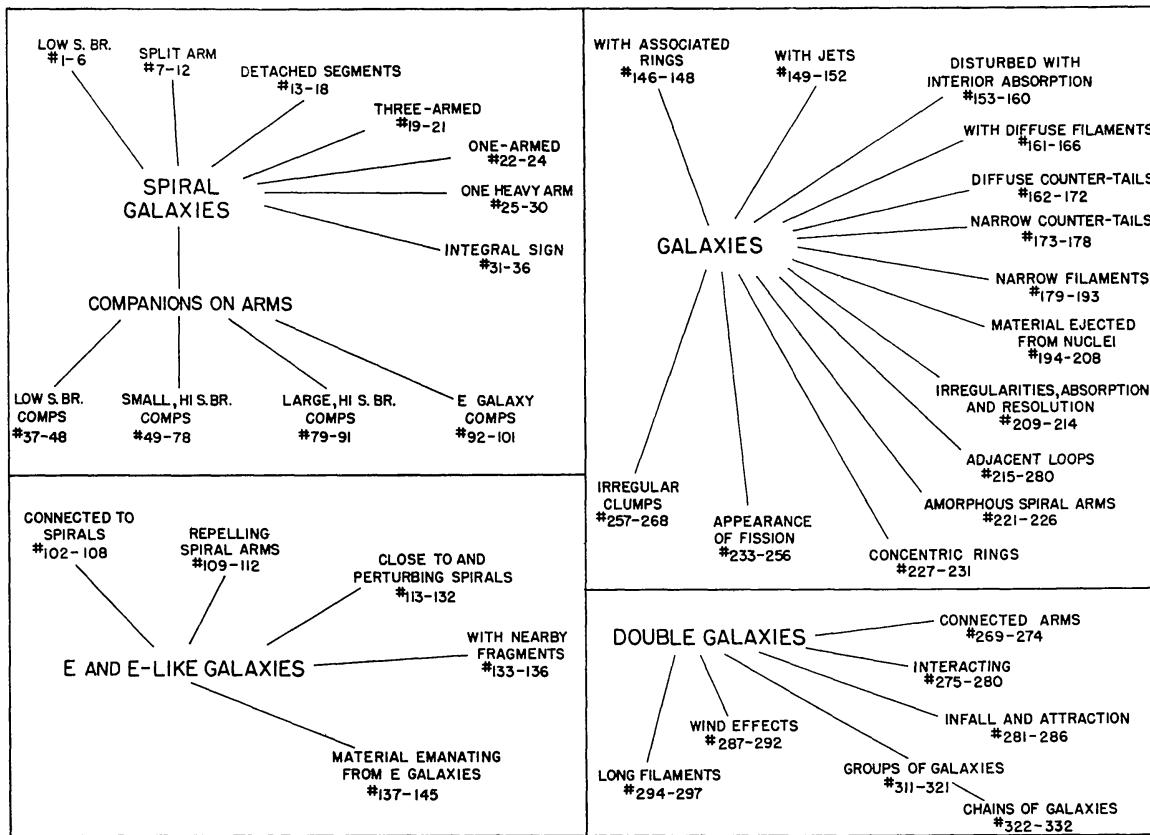


FIG. 1.—Plan of arrangements for the objects in the *Atlas*. The diagram shows major peculiarities which determine classification. Comments on additional peculiarities are given in Table 1. Characterization of peculiarities is sometimes descriptive rather than literal.